

On Relativistic Fluid Space Time Admitting Heat Flux of a Generalized Recurrent and Ricci-recurrent Kenmotsu Manifold*

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Abstract: The aim of the present paper is to study generalized recurrent and ricci-recurrent Kenmotsu manifold for relativistic fluid space time admitting heat flux with the time like generator U as the velocity vector field of the fluid and generator V as the heat flux vector field.

Keywords: Generalized recurrent Kenmotsu manifold, relativistic fluid space time, heat flux vector field

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1. Introduction

A Riemannian manifold (M_n, g) is called generalized recurrent¹ if its curvature tensor R satisfies the condition

$$(1) \quad (\nabla_X R)(Y, Z)W = \alpha(X)R(Y, Z)W + \beta(X)[g(Z, W)Y - g(Y, W)Z],$$

where α and β are two 1-forms, β is non- zero and these are defined by

$$(2) \quad \alpha(X) = g(X, A), \beta(X) = g(X, B),$$

where A, B are vector fields associated with 1-forms α and β respectively.

A Riemannian manifold (M_n, g) is called generalized Ricci-recurrent¹ if its Ricci-tensor S satisfies the condition

$$(3) \quad (\nabla_X S)(Y, Z) = \alpha(X)S(Y, Z) + (n-1)\beta(X)g(Y, Z),$$

where α and β are defined as in (2).

2. Kenmotsu Manifold

Let M be an almost contact manifold² equipped with an almost contact metric structure (ϕ, ξ, η, g) consisting of a $(1, 1)$ tensor field ϕ , a vector

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field ξ , a 1-form η and a compatible Riemannian metric g satisfying

$$(4) \quad \phi^2 = -I + \eta \otimes \xi, \quad \eta(\xi) = 1, \quad \phi(\xi) = 0, \quad \eta \circ \phi = 0,$$

$$(5) \quad g(X, Y) = g(\phi X, \phi Y) + \eta(X)\eta(Y),$$

$$(6) \quad g(X, \phi Y) = -g(\phi X, Y), \quad g(X, \xi) = \eta(X), \quad \text{for all } X, Y \in \chi(M).$$

An almost contact metric manifold M is called a kenmotsu manifold if it satisfies³

$$(7) \quad (\nabla_X \phi)Y = g(\phi X, Y)\xi - \eta(Y)\phi X, \quad X, Y \in \chi(M),$$

where ∇ is Levi-Civita connection of the Riemannian metric g . From the above equation it follows that

$$(8) \quad \nabla_X \xi = X - \eta(X)\xi,$$

$$(9) \quad (\nabla_X \eta)Y = g(X, Y)\xi - \eta(X)\eta(Y).$$

Moreover, the curvature tensor R and Ricci tensor S satisfy³

$$(10) \quad R(\xi, X)Y = \eta(Y)X - g(X, Y)\xi$$

and

$$(11) \quad S(X, \xi) = (1 - n)\eta(X), \quad \text{where } n = 2m + 1.$$

In 1980, D. ray⁴ defined the energy momentum tensor representing the matter distribution of a fluid space time admitting heat flux as

$$(12) \quad T(X, Y) = (\rho + \sigma)A(X)A(Y) + \sigma g(X, Y) + A(X)B(Y) + A(Y)B(X),$$

where ρ and σ are the density and isotropic pressure of the fluid, U is the unit time vector field such that $g(X, U) = A(X)$ and V is the heat flux vector field such that $g(X, V) = B(X)$. In 1983, B.O'Neil⁵ defined the form of Einstein's equation without cosmological constant as

$$(13) \quad S(X, Y) - \frac{1}{2}rg(X, Y) = kT(X, Y),$$

where k is the gravitational constant and r being the scalar curvature of the manifold.

3. Theorems

In this paper, we establish the following:

Theorem: *If a fluid space time of general relativity admitting heat flux is a generalized Ricci- recurrent Kenmotsu manifold with the time like generator U as the velocity vector field of the fluid and the generator V as the heat flux vector field, then*

$$r=2(1-n) \quad , \quad \text{provided } k=0.$$

Proof: In view of (12), (13) can be written as

$$S(X, Y) = \frac{1}{2}rg(X, Y) + k[(\rho + \sigma)A(X)A(Y) + \sigma g(X, Y) \\ + \{A(X)B(Y) + A(Y)B(X)\}],$$

or

$$(14) \quad S(X, Y) = \left(\frac{1}{2}r + k\sigma\right)g(X, Y) + k(\sigma + \rho)A(X)B(Y) \\ + k[A(X)B(Y) + A(Y)B(X)].$$

Assume that (M_n, g) be a generalized recurrent Kenmotsu manifold, then the Ricci-tensor S of (M_n, g) satisfies the condition (3) for all vector fields X, Y, Z . Taking $Z = \xi$ in (3) we have

$$(15) \quad (\nabla_X S)(Y, \xi) = (1-n)[\alpha(X) - \beta(X)]\eta(Y).$$

By the definition of covariant derivative of S we have

$$(16) \quad (\nabla_X S)(Y, \xi) = \nabla_X S(Y, \xi) - S(\nabla_X Y, \xi) - S(Y, \nabla_X \xi),$$

which in view of (8) and (11), can be written as

$$(17) \quad (\nabla_X S)(Y, \xi) = (1-n)\nabla_X \eta(Y) - (1-n)\eta(\nabla_X Y) - S(X, Y) \\ + (1-n)\eta(X)\eta(Y).$$

So using (9) we get

$$(18) \quad (\nabla_X S)(Y, \xi) = (1-n)g(X, Y) - S(X, Y).$$

From the equality of left hand side of the equation (15) and (18) we can write

$$(19) \quad (1-n)[\alpha(X) - \beta(X)]\eta(Y) = (1-n)g(X, Y) - S(X, Y).$$

Hence taking $Y = \xi$ in (19) and using (11), we obtain

$$(20) \quad \alpha(X) = \beta(X), \quad \text{for any vector field } X.$$

Using (20) in (19), we have

$$(21) \quad S(X, Y) = (1-n)g(X, Y).$$

From (21) and (14), we have $\frac{r}{2} = (1-n)$, provided $k = 0$. This implies $r = 2(1-n)$, provided $k = 0$. Hence the theorem.

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